Smadar Naoz Einstein Fellow, Harvard Smithsonian Center for Astrophysics ITC

Naoz & Narayan 2013, PRL

Plasma Meeting October 2013

 Coherent magnetic fields in galaxies and clusters ~ 10⁻⁶ Gauss e.g., Widrow 2002

47 16 **DECLINATION (J200**

M51 B~30µG

 Coherent magnetic fields in galaxies and clusters ~ 10⁻⁶ Gauss e.g., Widrow 2002

The dynamo mechanism?



e.g., Parker 1955; Ruzmaikin et al 1988; Kulsrud 1999; Brandenburg and Subramanian 2005; Kulsrud and Zweibel 2008 M51 B~30µG



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TTTK

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Magnetogenesis

Magnetogenesis

Top down

Magnetogenesis

Top down



Magnetogenesis

Top down





Magnetogenesis

Top down









Magnetogenesis

Top down











Magnetogenesis

Top down















t~0.4Myr



t~0.4Myr















Magnetic fields on linear over-densities

re-ionization re-ionization



Generate magnetic field from zero B

Generate magnetic field from zero B through vortex like motion:

Generate magnetic field from zero B through vortex like motion:



Generate magnetic field from zero B through vortex like motion:



Generate magnetic field from zero B through vortex like motion Q: How do we generate vorticity?
Generate magnetic field from zero B through vortex like motion Q: How do we generate vorticity? A: Biermann 1950: $\nabla n \times \nabla P \neq 0$

Generate magnetic field from zero B through vortex like motion Q: How do we generate vorticity? $\nabla n \times \nabla P \neq 0$ (No Plasma:)

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P/n



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P/n (T/n)





Generate magnetic field from zero B through vortex like motion Q: How do we generate vorticity?

 $\nabla n \times \nabla P \neq 0$ (No Plasma:)





Where do we find vorticity in the dark ages?
Where is the plasma?

P/n (T/n) $\nabla n \times \nabla P \neq 0$

2. Where is the plasma?

P/n (T/n) $\nabla n \times \nabla P \neq 0$

2. Where is the plasma?

Linear T fluc. / gas density fluc.

P/n (T/n) $\nabla n \times \nabla P \neq 0$













Where do we find vorticity in the dark ages?
Where is the plasma?

COMPTON SCATTERING



COMPTON SCATTERING



+ THE STREAM VELOCITY

Tseliakhovich & Hirata 2010







+ THE STREAM VELOCITY

Tseliakhovich & Hirata 2010

- |v_b-v_{dm}|≈30 km/sec at Recombination time =Mach 5
- scales as 1/a



e.g.,effects on structure formation Greif et al 2011, Stacy et al 2011, Naoz et al 2012,2013, Visbal et al 2012

COMPTON SCATTERING



+ THE STREAM VELOCITY

Tseliakhovich & Hirata 2010

- |v_b-v_{dm}|≈30 km/sec at Recombination time =Mach 5
- scales as 1/a



0.8 1σ stream vel 0.6 z = 30^q9/^L 0.4 z=100 e.g., effects on structure formation 0.2 Greif et al 2011, Stacy et al 2011, Naoz et al 2012,2013, 10 10^{2} 103 10^{4} Visbal et al 2012 $k [Mpc^{-1}]$ Naoz & Narayan 2013, PRL

COMPTON SCATTERING











THE EFFECTS OF REIONIZATION

 $\dot{B} \sim \overline{T} \nabla \delta_{e} \times \nabla \delta_{T}$

- T→10⁴ K
- • $\delta_e \rightarrow$ Larger in 10-10²
- • δ_T → Larger in 10-10²

THE EFFECTS OF REIONIZATION

 $\dot{B} \sim \overline{T} \nabla \delta_{\rho} \times \nabla \delta_{\tau}$

- T→10⁴ K
- • $\delta_e \rightarrow$ Larger in 10-10²
- • δ_T → Larger in 10-10²

Pre-reionization $10^{-25} \cdot 10^{-24} \text{ G}$ B- 310^{-18} G present day











PRIMORDIAL MAGNETIC FIELDS



PRIMORDIAL MAGNETIC FIELDS

