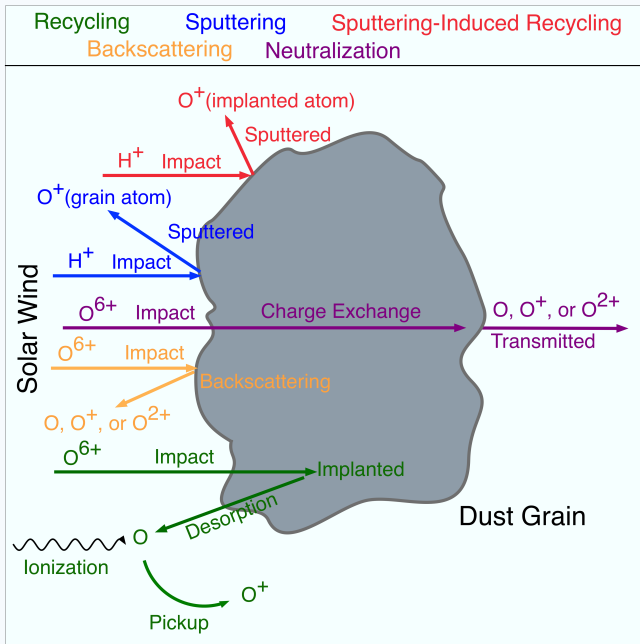
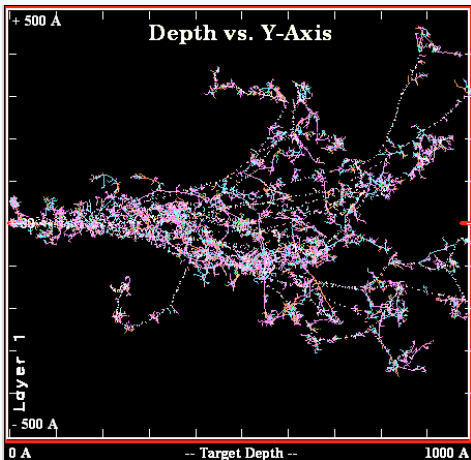


**INNER SOURCE  $C^+$  /  $O^+$  PICKUP IONS PRODUCED BY SOLAR WIND RECYCLING,  
NEUTRALIZATION, BACKSCATTERING, SPUTTERING, AND SPUTTERING-INDUCED  
RECYCLING**

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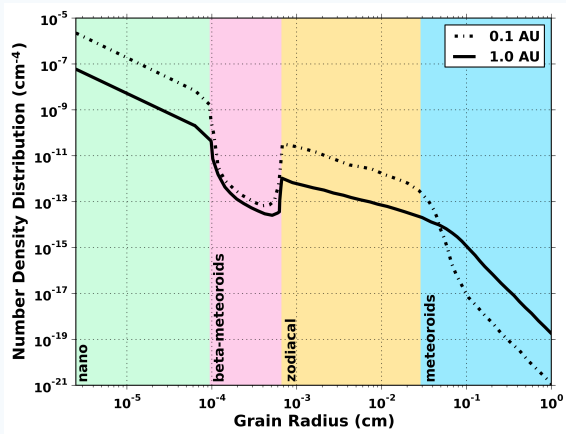
P. R. Quinn, N. A. Schwadron, E. Möbius, A. Taut, L. Berger





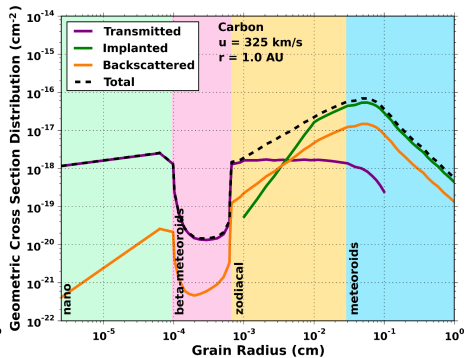
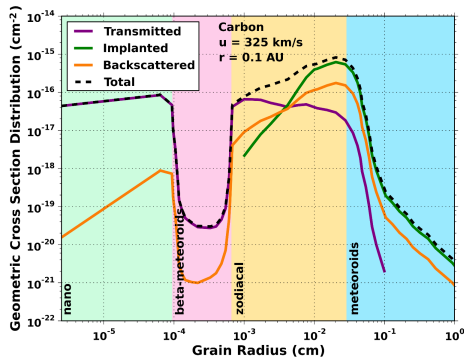
- C = 2.39
  - O = 3.98
  - Na = 0.056
  - Mg = 1.015
  - Al = 0.070
  - S = 0.417
  - Ca = 0.047
  - Cr = 0.016
  - Fe = 0.705
  - Ni = 0.024
- Average chondritic IDP composition relative to Si.
  - Consider fluffy IDPs with equivalent density of  $0.001 \text{ g/cm}^3$ .
  - Use SRIM to simulate solar wind ions through grains of varying size to determine the percent of ions transmitted, implanted, backscattered, and the sputtering yield.

# GRAIN SIZE DISTRIBUTION



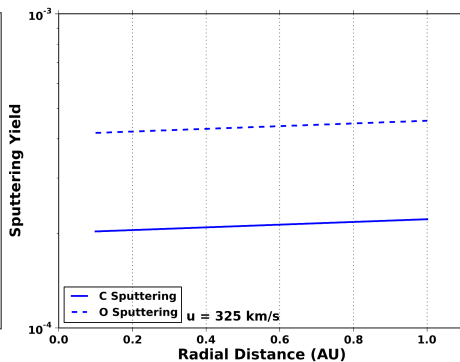
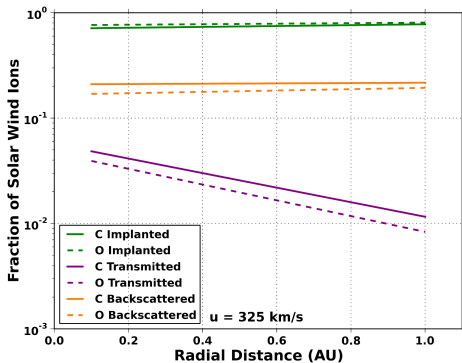
- Grains spiral toward the sun due to the Poynting-Robertson effect.
- Larger grains break apart into smaller grains due to dust-dust collisions.

# SRIM RESULTS AND GRAIN GEOMETRIC CROSS SECTION DISTRIBUTION



•  $\Gamma(r, s) = n(r, s)\pi s^2.$

# FRACTION OF PUIs PRODUCED - RADIAL DEPENDENCE



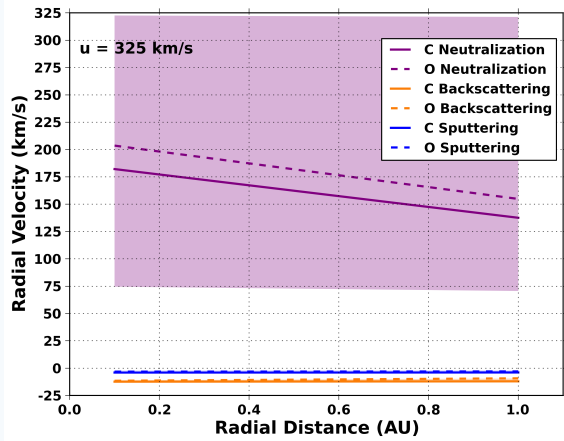
- Integrate over all grain sizes to get the total percent of PUIs produced for each mechanism.

- $k_{N,R,B}(r, u) = \int k_{N,R,B}(r, s, u) n(r, s) \pi s^2 ds$

- Similarly for the sputtering yield.

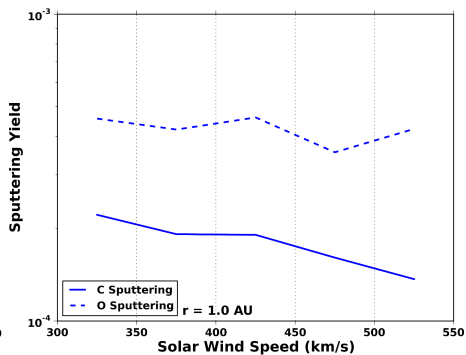
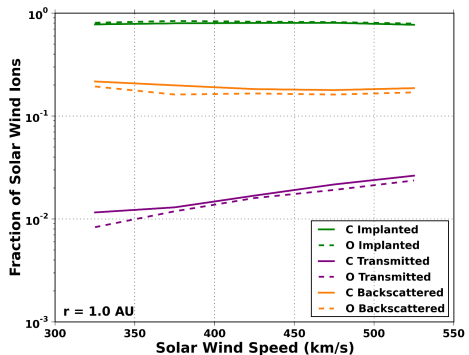
- $Y(r, u) = \int Y(r, s, u) n(r, s) \pi s^2 ds$

## AVERAGE RADIAL VELOCITY



- Average radial velocity of PUIs weighted by  $k_{N,B}(r, u, s)$  or  $Y(r, u)$  at each grain size.
- Shaded area is  $1\sigma$  about the average radial velocity (very small for sputtering and backscattering).
- The velocity of recycled ions is that of the grain.

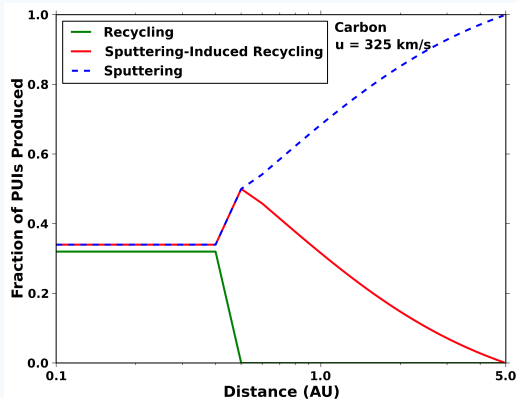
## SOLAR WIND SPEED DEPENDENCE



- Repeat entire process for solar wind speeds from 325 km/s to 525 km/s.
- Now we have the following as a function of distance from the sun and solar wind speed:
  - Fraction of PUIs transmitted (neutralization), implanted (recycling and sputtering-induced recycling), and backscattering,  $k_{N,R,B}(r, u)$
  - Sputtering yield,  $Y(r, u)$ .
  - Average radial velocity of ions after production,  $v_{N,R,B,S}(r, u)$ .



# SATURATION, DISSOCIATION, AND SPUTTERING-INDUCED RECYCLING



- Saturation

- Saturation occurs when a one-to-one correlation between implanted solar wind ions and dust grain atoms is reached.
- Solar wind flux close to the sun is high enough to saturate grains during their Poynting-Robertson lifetime.

- Dissociation

- Implanted C and O form compounds with implanted H.
- Assume that C is desorbed as CH and O is desorbed as OH.
- $\text{CH} \rightarrow \text{C} + \text{H}$  or  $\text{CH} \rightarrow \text{CH}^+ + \text{e}^-$ .

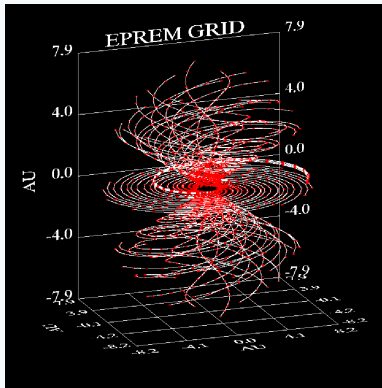
- Sputtering-Induced Recycling

- Once ions are implanted, they have a chance of being sputtered.
- Within the saturation zone, there's 50% of sputtering a grain atom or implanted ion.

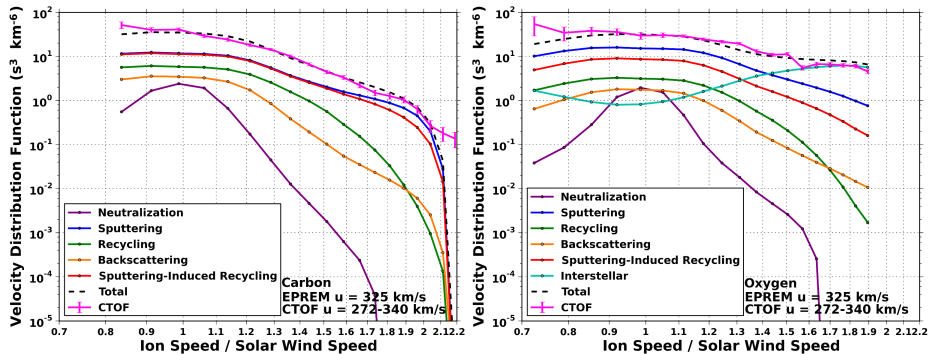
- Fraction of PUIs produced by recycling, sputtering, and sputtering-induced recycling,  $\kappa_{R,S,SIR}(r, u)$

# PRODUCTION RATE PER UNIT VOLUME

- $S_{\text{R}}^{+}(r, u) = \kappa_{\text{R}}(r, u)k_{\text{R}}(r, u)S_{\text{impact}}(r)P^{+}(r) \left[ \left( \frac{\beta_{\text{diss}}}{\beta_{\text{diss}} + \beta_{\text{ion}}} \right) \left( \frac{\beta_{\text{ion}}^0}{\beta_{\text{ion}}^0 + \beta_{\text{ion}}^{+}} \right) + \left( \frac{\beta_{\text{ion}}}{\beta_{\text{diss}} + \beta_{\text{ion}}} P_{\text{diss}}(r) \right) \right]$
- $S_{\text{N}}^{+}(r, u) = k_{\text{N}}(r, u)S_{\text{impact}}(r) [\eta^{+}P^{+}(r) + \eta^0P^0(r)]$
- $S_{\text{B}}^{+}(r, u) = k_{\text{B}}(r, u)S_{\text{impact}}(r) [\eta^{+}P^{+}(r) + \eta^0P^0(r)]$
- $S_{\text{S}}^{+}(r, u) = \kappa_{\text{S}}(r, u)Y(r, u)S_{\text{impact}}P^{+}(r)$
- $S_{\text{SIR}}^{+}(r, u) = \kappa_{\text{SIR}}(r, u)Y(r, u)S_{\text{impact}}P^{+}(r) \left[ \left( \frac{\beta_{\text{diss}}}{\beta_{\text{diss}} + \beta_{\text{ion}}} \right) \left( \frac{\beta_{\text{ion}}^0}{\beta_{\text{ion}}^0 + \beta_{\text{ion}}^{+}} \right) + \left( \frac{\beta_{\text{ion}}}{\beta_{\text{diss}} + \beta_{\text{ion}}} P_{\text{diss}}(r) \right) \right]$
- $S_{\text{I}}^{+}(r, \theta) = \beta_{\text{ion}}^0 \left( \frac{r_1}{r} \right)^2 n_{\text{O}}(r, \theta)P^{+}(r)$

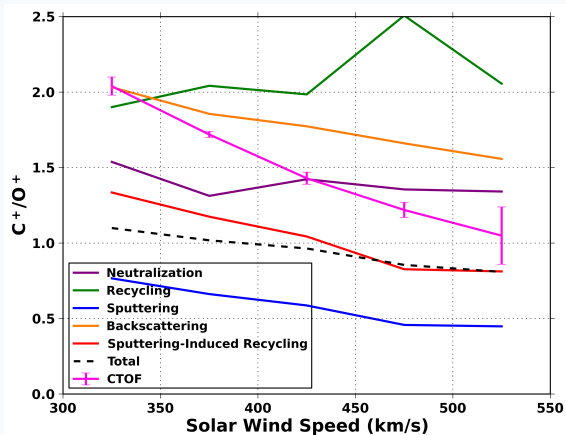


# VDF COMPARISON TO CTOF



- Compare EPREM to SOHO/CTOF at 1 AU.
- Method from Taut et al. 2015 is used to derive normalized count rates of  $C^+$  and  $O^+$  as a function of  $w = v/u$ .
- **New constraint - broad VDF at 1 AU with possible cutoff near  $w = 2$ .** Implies PUIs near 1 AU are produced with near zero speed.

# $C^+/O^+$ SOLAR WIND DEPENDENCE



- Recycling  $C^+/O^+ > 1$  due to dissociation rate of CH being much higher than OH.
- Sputtering-Induced Recycling  $C^+/O^+ > 1$  for slow solar wind speeds but is just below 1 for higher speeds.
- Total  $C^+/O^+ > 1$  for slow solar wind speeds and now decreases with speed.

# CONSTRAINTS

Constraint	Recycling	Neutralization	Backscattering	Sputtering	Sputtering-Induced Recycling
Broad VDF at 1 AU with cutoff near $w = 2$	No	Possibly	Yes	Yes	Yes
$C^+/O^+ > 1$	Yes	Yes	Yes	No	Mostly
Decreasing $C^+/O^+$ with $u$	No	Yes	Yes	Yes	Yes
Solar wind-like composition	Yes	Yes	Yes	No	Yes
Stability over the solar cycle	Yes	Unlikely	Yes	Yes	Yes

- Backscattering and SIR satisfy the most constraints, but the total  $C^+/O^+$  ratio does not match CTOF due to the over-production of  $O^+$ .
- Possibility that silicate IDPs break down. For example  $MgSiO_3 \rightarrow MgO + SiO_2$  where  $MgO$  has higher survivability near the sun.
- We test the production of sputtering from  $MgO$  IDPs. Results show a 30-40% decrease in  $O^+$  production – making the total  $C^+/O^+$  ratio about 1.37-0.95 for 325-525 km/s solar wind speed. These results move toward better agreement with CTOF.
- This suggests that the IDP composition may indeed change inside 1 AU.

- Determined a new constraint on the inner source PUI production mechanism: a broad VDF at 1 AU with possible cutoff near  $w = 2$  which implies PUIs near 1 AU are injected into the solar wind at near zero speeds.
- Introduced two new mechanisms for inner source PUI production: backscattering and sputtering-induced recycling.
- Backscattering and sputtering-induced recycling satisfy the most constraints.
- The dominant mechanisms (based on intensity) are sputtering and sputtering-induced recycling.
- The composition of our simulated PUIs does not perfectly agree with CTOF observations most likely due to the high production of sputtered  $O^+$ . We showed that a change in IDP composition may reduce the amount of sputtered  $O^+$  and agrees better with CTOF. Observational evidence of PUI composition and dust grain composition inside 1 AU is necessary to investigate this idea further.